

## Measurement of $^{25}\text{Al}+p$ resonant elastic scattering relevant to the astrophysical reaction $^{22}\text{Mg}(\alpha, p)^{25}\text{Al}$

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Type-I X-ray bursts are the most frequently observed thermonuclear explosions in nature.<sup>1)</sup> They occur on the surface of accreting neutron stars in low-mass X-ray binary systems. The investigation of X-ray bursts can help us understand the properties of a neutron star and the underlying physics.

The bursts are driven by the tripe- $\alpha$  reaction, the  $\alpha p$ -process, and the  $rp$ -process. After breakout from the hot CNO cycle, the nucleosynthesis path is characterized by the  $\alpha p$ -process.<sup>2)</sup> The  $\alpha p$ -process is a sequence of  $\alpha$ - and proton-induced reactions that transport nuclear material from the CNO cycle toward the region of heavier proton-rich nuclei.

The X-ray light curve is the main direct observable of X-ray bursts and could be affected significantly by the  $\alpha p$ -process. According to a recent sensitivity study by Cyburt,<sup>3)</sup> the  $^{22}\text{Mg}(\alpha, p)^{25}\text{Al}$  reaction is thought to be the most sensitive one during the  $\alpha p$ -process and may have a prominent impact on the burst light curve. However, with scarce experimental information on this reaction, the reaction rate in the calculations of X-ray bursts is estimated based on statistical models.

A measurement of  $^{25}\text{Al}+p$  resonant elastic scattering has been performed to experimentally examine the  $^{22}\text{Mg}(\alpha, p)^{25}\text{Al}$  reaction rates. The experiment was performed using the CNS radioactive ion beam separator (CRIB),<sup>4)</sup> installed by the Center for Nuclear Study (CNS), University of Tokyo, in the RIKEN Accelerator Research Facility. A primary beam of  $^{24}\text{Mg}^{8+}$  was accelerated up to 8.0 MeV/nucleon by the AVF cyclotron with an average intensity of 1 e $\mu$ A. The primary beam bombarded a liquid-nitrogen-cooled D<sub>2</sub> gas target to produce a secondary  $^{25}\text{Al}$  beam via the  $^{24}\text{Mg}(d, n)^{25}\text{Al}$  reaction in inverse kinematics. The  $^{25}\text{Al}$  beam was separated by the CRIB separator. The  $^{25}\text{Al}$  beam, with an average energy of 5.68 MeV/nucleon and an average intensity of  $2.0 \times 10^5$  pps, was then delivered to the F3 experimental chamber, where it bom-

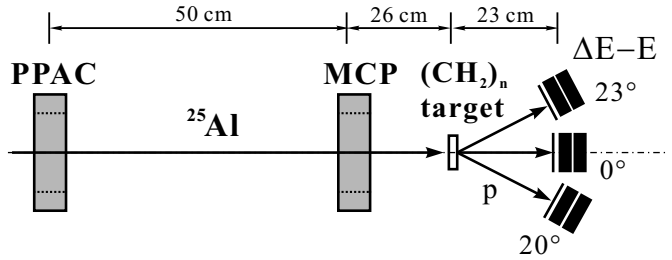


Fig. 1. Schematic diagram (top view) of the experimental setup at the F3 chamber.

barded a 13.95-mg/cm<sup>2</sup>-thick  $(\text{CH}_2)_n$  target and a 18.02-mg/cm<sup>2</sup>-thick C target in which the beam was stopped. The C target was used to evaluate the background contributions. After passing through a Wien filter, the  $^{25}\text{Al}$  beam purity can be up to 80%.

The setup at the F3 experimental chamber is shown in Fig. 1. A parallel-plate avalanche counter (PPAC) and a micro channel plate (MCP) were used for measuring the time and position information of the beam particles. The beam particles were identified in an event-by-event mode using the abscissa of MCP and the time of flight between MCP and the RF signal provided by the cyclotron.

The recoiling light particles from the  $^{25}\text{Al}+p$  reaction were measured using three sets of Si telescopes at average angles of  $\theta_{lab} \approx 0^\circ$ ,  $20^\circ$ , and  $23^\circ$ , respectively. Each telescope consisted of a 65- $\mu\text{m}$ -thick double-side-strip ( $16 \times 16$  strips) silicon detector and two 1500- $\mu\text{m}$ -thick pad detectors. The recoiling particles were clearly identified using the  $\Delta E$ -E method. An array of ten NaI(Tl) detectors was mounted directly above the target and used to detect the  $\gamma$  rays from the decay of the excited states in  $^{25}\text{Al}$ . The data analysis is in progress.

### References

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